

# ON THE DERIVATION OF THE EQUATIONS OF ANISOTROPIC MAGNETOHYDRODYNAMICS

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*PMM Vol. 26, No. 6, 1962, pp. 1092-1093*

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*(Received September 13, 1962)*

In the case of strong magnetic fields, the viscous stress tensor and the heat flow vector in magnetohydrodynamics depend on the magnetic field.

A two-fluid study of a fully ionized gas was carried out in [1], where the viscous stress tensors and the heat flow vectors of the ions and the electrons are calculated separately. Moreover, it is assumed that the temperature  $T_1$  of the electron gas differs from  $T_2$  of the ions.

In [2], based on the results of [1], expressions are obtained for the components of the viscous stress tensor and the heat flow vector in the one-fluid approximation. It is also assumed that

$$T_1 = T_2 = T \text{ and } \sqrt{m_1} \ll \sqrt{m_2} \quad (m_1 \text{ and } m_2 \text{ are the masses of the electrons and ions)}$$

However, [2] contains certain errors, which, if uncorrected, make the resulting equations inapplicable. In particular, it is necessary to recompute the transfer coefficients in the expressions for the components of the heat-flow vector. Furthermore, corrections were made of some misprints contained in [1], for which the author is indebted to a communication from S.I. Braginskii.

In expression (3.18) for  $q_u$  in [1], instead of the minus sign in front of the curly brackets, there is a plus sign. In formulas (4.14), instead of  $b''$ , read  $-b''$ . In [2], in describing the viscous stress tensor, errors made by Chapman and Cowling [3] and corrected in [4], were repeated in the signs of certain terms.

Using the same notation as in [2], we write the expressions for the components of the viscous stress tensor in the one-fluid approximation (furthermore, we consider only fully ionized gases consisting of

electrons and monatomic ions)

$$\begin{aligned}
 \pi_{zz} &= -\eta e_{zz} \\
 \pi_{xx} &= -\eta \left\{ b'_2 e_{xx} + \frac{1}{2} (b'_2 - 1) e_{zz} + 2\omega_2 \tau_2 b''_2 e_{xy} \right\} \\
 \pi_{yy} &= -\eta \left\{ b'_2 e_{yy} + \frac{1}{2} (b'_2 - 1) e_{zz} - 2\omega_2 \tau_2 b''_2 e_{xy} \right\} \\
 \pi_{xy} = \pi_{yx} &= -\eta \{ b'_2 e_{xy} - b''_2 \omega_2 \tau_2 (e_{xx} - e_{yy}) \} \\
 \pi_{xz} = \pi_{zx} &= -\eta \{ b'_1 e_{xz} + b''_1 \omega_2 \tau_2 e_{yz} \} \\
 \pi_{yz} = \pi_{zy} &= -\eta \{ b'_1 e_{yz} - b''_1 \omega_2 \tau_2 e_{xz} \}
 \end{aligned} \tag{1}$$

The values of the coefficients here are the same as in [2]. To obtain the expressions for the components of the heat flow vector, we use the formula in [2]

$$\begin{aligned}
 q_i &= \sum_s q_{si} + \frac{5}{2} \sum_s n_s k T c_{si} + \sum_{s, k} \pi_{sik} c_{sk} + \frac{1}{2} \sum_s n_s m_s c_{si} c_s^2 \\
 c_s &= v_s - v \quad (s = 1, 2)
 \end{aligned}$$

Since  $c_s$  is proportional to the current density, then, neglecting the last term as a quadratic term in  $j$ , we also neglect the third term on the right by comparing it with the second. This is permissible for a continuous medium, when the characteristic frequency  $\Omega$  is much smaller than the "collision frequency" of the charged particles, if, using some estimates of [5], we assume that  $(\Omega/\omega_2)\tau_1/t$  is not greater than unity in order of magnitude. ( $t$  is the characteristic time of the problem.) Moreover, it is assumed that  $v = v_2$  (this being true for  $m_1 v_1 \ll m_2 v_2$ ).

Utilizing these corrected results of [1] we may obtain for the components of the heat flow vector in the one-fluid approximation, to first order accuracy in the current density

$$\begin{aligned}
 q_z &= -\lambda \left( \frac{\partial T}{\partial z} + v j_z \right), & q_x &= -\lambda \left( \kappa \frac{\partial T}{\partial x} - \omega_1 \tau_1 \kappa' \frac{\partial T}{\partial y} + v' j_x - \omega_1 \tau_1 v' j_y \right) \\
 & & q_y &= -\lambda \left( \kappa \frac{\partial T}{\partial y} + \omega_1 \tau_1 \kappa' \frac{\partial T}{\partial x} + v' j_y + \omega_1 \tau_1 v' j_x \right)
 \end{aligned} \tag{2}$$

Here

$$\begin{aligned}
 \lambda &= 1.58 \frac{pk\tau_1}{m_1}, & \kappa &= (1.47\omega_1^2\tau_1^2 + 3.77) \Delta_1 + \frac{\omega_2\tau_2}{\omega_1\tau_1} (0.633\omega_2^2\tau_2^2 + 0.837) \Delta_2 \\
 \omega_1 &= \frac{eH}{m_1c} > 0, & \kappa' &= (0.791\omega_1^2\tau_1^2 + 6.88) \Delta_1 - \left( \frac{\omega_2\tau_2}{\omega_1\tau_1} \right)^2 (0.791\omega_1^2\tau_1^2 + 1.47) \Delta_2 \\
 \iota &= 2.03 \frac{TH}{\omega_1\tau_1 pc}, & \iota' &= (1.58\omega_1^4\tau_1^4 + 26.6\omega_1^2\tau_1^2 + 7.66) \frac{TH\Delta_1}{\omega_1\tau_1 pc}
 \end{aligned} \tag{3}$$

$$v'' = (0.949\omega_1^2\tau_1^2 + 1.93) \frac{TH\Delta_1}{\omega_1\tau_1\rho c} \quad (|e_1| = e)$$

$$\frac{1}{\Delta_1} = (\omega_1^4\tau_1^4 + 14.79\omega_1^2\tau_1^2 + 3.77), \quad \frac{1}{\Delta_2} = (\omega_2^4\tau_2^4 + 2.70\omega_2^2\tau_2^2 + 0.677)$$

We observe that, aside from correcting some signs and values of the coefficients, the nature of the dependence of  $v''$  on  $\omega_1\tau_1$  obtained here is different from that in [2], and for large values of  $\omega_1\tau_1$ , it agrees with the results of [4] to within the coefficients. We recall that the magnetic field  $H$  is directed along the  $z$ -axis.

Substituting (1) and (2) into the energy equation

$$\rho c_v \frac{dT}{dt} = -p \operatorname{div} \mathbf{v} - \pi_{\alpha\beta} \nabla_\alpha v_\beta - \nabla \mathbf{q} + \mathbf{j} \left( \mathbf{E} + \frac{1}{c} \mathbf{v} \times \mathbf{H} \right)$$

we finally obtain

$$\begin{aligned} \rho c_v \frac{dT}{dt} = & -p \operatorname{div} \mathbf{v} + \mathbf{j} \left( \mathbf{E} + \frac{1}{c} \mathbf{v} \times \mathbf{H} \right) + \eta b_2' \left[ 2 \sum_i \left( \frac{\partial v_i}{\partial x_i} \right)^2 - (\operatorname{div} \mathbf{v})^2 \right] + \\ & + \eta b_2' \left( \frac{\partial v_x}{\partial y} + \frac{\partial v_y}{\partial x} \right)^2 + \eta b_1' \left[ \left( \frac{\partial v_x}{\partial z} + \frac{\partial v_z}{\partial x} \right)^2 + \left( \frac{\partial v_y}{\partial z} + \frac{\partial v_z}{\partial y} \right)^2 \right] + \frac{1}{3} \eta (\operatorname{div} \mathbf{v})^2 + \\ & + \eta (b_2' - 1) \frac{\partial v_z}{\partial z} \left( 2 \operatorname{div} \mathbf{v} - 3 \frac{\partial v_z}{\partial z} \right) + \operatorname{div} (\lambda \kappa \nabla T) + \frac{\partial}{\partial z} \lambda (1 - \kappa) \frac{\partial T}{\partial z} + \operatorname{div} (\lambda \nu' \mathbf{j}) + \\ & + \frac{\partial}{\partial z} \lambda (1 - \nu') j_z - \frac{\partial}{\partial x} (\lambda \omega_1 \tau_1 \kappa' \frac{\partial T}{\partial y}) + \frac{\partial}{\partial y} (\lambda \omega_1 \tau_1 \kappa' \frac{\partial T}{\partial x}) - \frac{\partial}{\partial x} (\lambda \rho_1 \tau_1 \nu' j_y) + \frac{\partial}{\partial y} (\lambda \omega_1 \tau_1 \nu' j_x) \end{aligned} \quad (4)$$

If the currents are absent, and the magnetic field  $\mathbf{H} = 0$ , then equation (4) becomes the energy equation of ordinary hydrodynamics.

In addition to comparing with ordinary heat conduction when  $\mathbf{H} = \mathbf{j} = 0$ , the terms in (2) describe well-known physical phenomena (e.g. [6]): the terms with  $\nu$  and  $\nu'$  give Thomson's effect, those with  $\omega_1\tau_1\kappa'$  the Leduc-Rigi effect, and those with  $\omega_1\tau_1\nu''$  the Etingshausen effect. The last two effects are connected with the Larmor rotation of the electrons in the magnetic field.

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Translated by C.K.C.